

## f(R) THEORIES

These theories arise from a straightforward generalization of the Einstein-Hilbert action which becomes a general function of R

$$S_J = \frac{1}{2\kappa} \int d^4x \sqrt{-g} (\xi R - V(\xi) + 2\kappa \mathcal{L}_m) \quad \text{JORDAN FRAME}$$

INDEPENDENT DEGREES OF FREEDOM

$V(\xi)$  fixed by the specific functional form  $f(R)$

$$V(\xi) \equiv f(R(\xi)) - \xi R(\xi) \quad \text{where } \xi = \frac{df}{dR}$$

## BIANCHI I UNIVERSE

Homogeneous and anisotropic (flat  ${}^3R = 0$ ) model

$$ds^2 = N^2(t) dt^2 - \eta_{ab}(t) \omega^a \otimes \omega^b \quad \text{4D LINE ELEMENT}$$

$$\eta_{ab} = R_0^2 e^{\alpha} (e^{\beta})_{ab}$$

- CHANGE OF VOLUME
- CHANGE OF SHAPE

INITIAL RADIUS OF THE UNIVERSE

The matrix  $\beta_{ab}$  satisfies the condition  $\text{Tr} \beta = 0$ , ensuring the volume depends only on the conformal factor  $\alpha$

$$\mathcal{V}_{Uni} \sim R_0^3 e^{3\alpha}$$

# ON THE EMERGENCE OF A CLASSICAL UNIVERSE FROM A QUANTUM f(R) COSMOLOGY IN THE JORDAN FRAME

Bianchi I Universe in the Jordan frame of an  $f(R)$  theory is associated to the reduced ADM action

$$S_B = \int d\xi \left( p_\alpha \frac{d\alpha}{d\xi} + p_+ \frac{d\beta_+}{d\xi} + p_- \frac{d\beta_-}{d\xi} - H_{ADM} \right)$$

TEMPORAL GAUGE

$$\dot{\xi} = 1 \longrightarrow N = \frac{3e^{3\alpha}}{2(\xi H_{ADM} - p_\alpha)}$$

The Hamiltonian constraint is classically solved (RPSQ)

QUANTISING  $H_{ADM}$  yields the Schrödinger equation

$$H_{ADM} \equiv \frac{1}{\xi} \left( p_\alpha + \sqrt{p_\alpha^2 - p_+^2 - p_-^2 - 6\xi e^{3\alpha} (V(\xi) + \Lambda)} \right)$$

$$i\partial_\xi \psi = -\frac{1}{\xi} \left( -i\partial_\alpha + \sqrt{-\partial_\alpha^2 + \partial_+^2 + \partial_-^2 - 6\xi e^{3\alpha} \Lambda} \right) \psi$$

• ASSUMPTION:  $\Lambda \gg V(\xi)$

• WITHOUT  $\Lambda$

$$\psi(\alpha, \beta_\pm, \xi) = e^{i(k_\alpha \alpha + k_+ \beta_+ + k_- \beta_-)} \xi^{i(k_\alpha + \sqrt{k_\alpha^2 - k_+^2 - k_-^2})}$$

Use a Gaussian distribution  $A(k_\alpha, k_+, k_-)$  to construct a localized wave packet

$$\Psi(\alpha, \beta_\pm, \xi) = \int_{-\infty}^{+\infty} dk_\alpha dk_+ dk_- A(k_\alpha, k_+, k_-) \psi(\alpha, \beta_\pm, \xi)$$

RESULT

The wave packet becomes localised (classicalisation) but is dynamically unstable  
(NOT TRUE FOR FLRW  $\rightarrow$  REACHES THE SINGULARITY IN A QUASI-CLASSICAL WAY)

## BIANCHI I VS FLRW

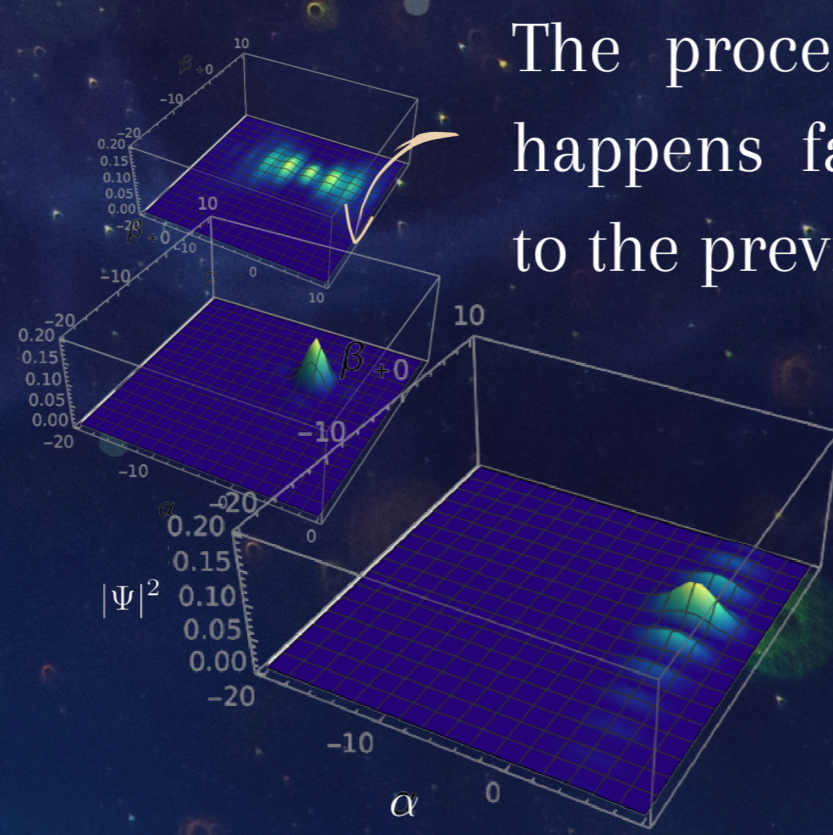
• WITH  $\Lambda$

$$\psi(\alpha, \xi) = e^{ik_\alpha \alpha} e^{i \left( \frac{3\xi \Lambda e^{3\alpha}}{k_\alpha} + 2ik_\alpha \log(\xi) \right)}$$

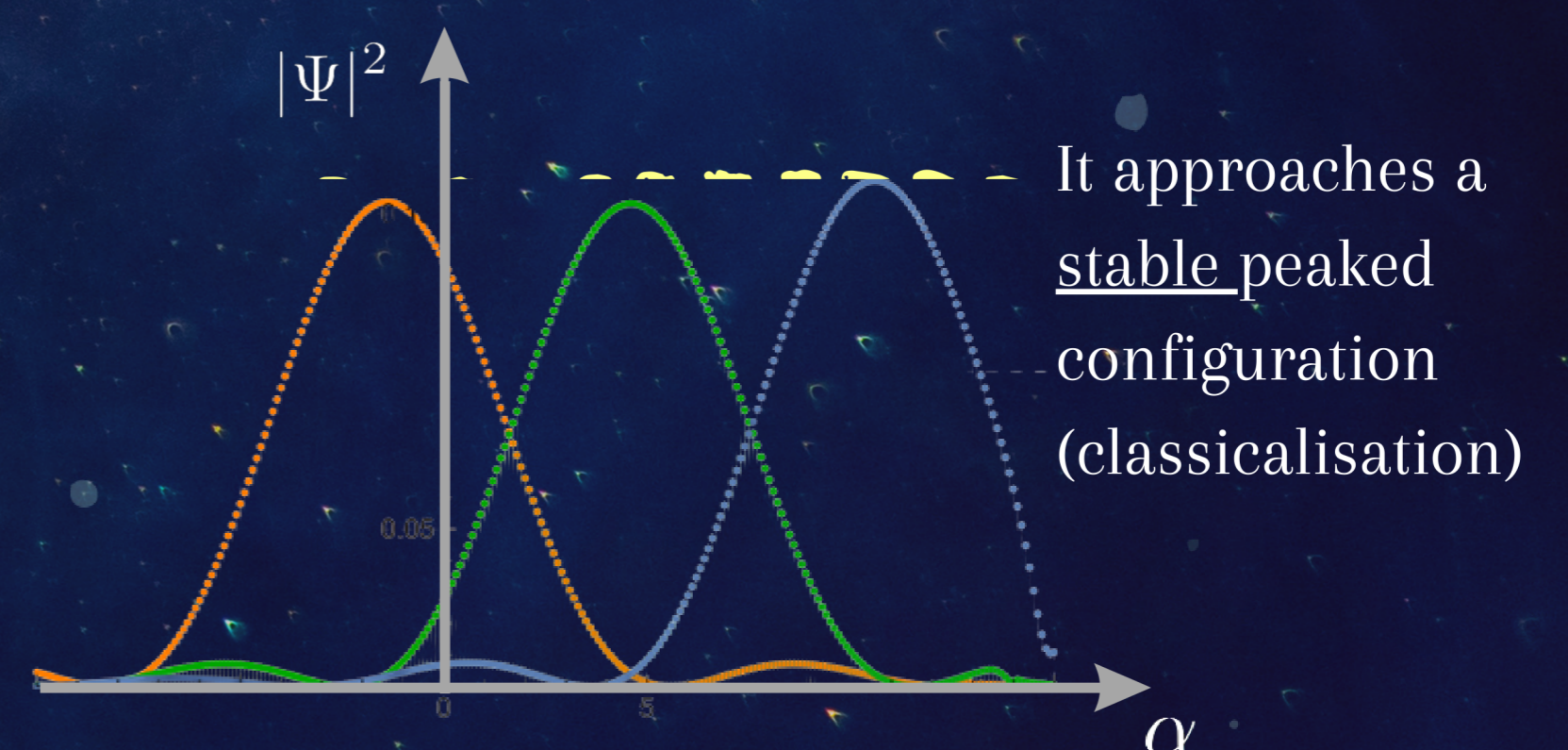
The introduction of  $\Lambda$  seems not to be the mechanism to make the model a classical one

- ASSUMPTION: On-set of the inflationary phase (and we neglect  $V(\xi)$ )

- TO INVESTIGATE: At late(r) times  $V(\xi)$  could play a significant role



The process of localisation happens faster with respect to the previous case!



It approaches a stable peaked configuration (classicalisation)

## REFERENCES

- MDA, G. Montani, *EPJC* **83**, 285 (2023)  
MDA, G. Montani, *Phys. Rev. D* **101**, 103532 (2020); MDA, L. Figurato, G. Montani, *Phys. Rev. D* **104**, 024054 (2021)

