# **Common challenges in quantum gravity**

**Bianca Dittrich**, **Perimeter Institute** 

Quantum Gravity 2023, Radboud University Nijmegen









## Diversity of ideas



Important to learn from each other, and to look for shared challenges and principles.

I will present a biased selection of challenges.

[de Boer et al: Frontiers of quantum gravity, shared challenges, converging directions (Snowmass 2021)]





### Our universe is Lorentzian

### ... most of the times.

# Euclidean Quantum Gravity Why???

Why did/do we do Euclidean quantum gravity?  $Z_E = \int \mathscr{D}geom_E \exp(-S(geom_E)) \quad V_s \quad Z_L = \int \mathscr{D}geom_L \exp(\imath S(geom_L))$ 

- Wick rotation: successful strategy in QFT
- Computational techniques: Monte Carlo vs. few techniques for complex amplitudes
- Thermodynamic interpretation

Euclidean Quantum Gravity



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• Lorentzian configuration space very different from Euclidean configuration space: No well-defined Wick rotation.



• Lattice simulation: tends to drive system into region where conformal factor is maximized (spikes, branched polymers, ...)

[Ambjorn, Loll, Jurkiewicz et al. 98+: Causal Dynamical Triangulations]

[Talks by Ambjorn, Goerlich]



[Goerlich]





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- Lattice simulation: tends to drive system into region where conformal factor is maximized (spikes, branched polymers, ...)

• Lorentzian configuration space very different from Euclidean configuration space: No well-defined Wick rotation.

• No destructive interference which might cancel out unwanted configurations

[Carlip, Carlip, Surya 22: Causal Sets]



[Ambjorn, Loll, Jurkiewicz et al. 98+: Causal Dynamical Triangulations]

[Talks by Ambjorn, Goerlich]



[Goerlich]

[Talks by Carlip, Surya]



## Computational techniques for complex amplitudes

- deformation of integration contour, possible choice Lefschetz (Semi-classical and numerical)
- Holomorphic gradient flow / Monte Carlo on Lefschetz thimb (Numerical, "experimental")
- Acceleration operators for series convergence (for sums and (Numerical)
- Tensor network renormalization (Numerical, so far up to three dimensional systems)
- Asymptotic Safety (Lorentzian configuration space, Lorentzian renormalization flo
- Machine learning, quantum simulations, .... (To be explored: Scaling?)

[Talks by Asante, Feldbrugge]

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Challenge: Methods for large scale systems.







## Acceleration techniques for series convergence

[Schmidt 41, Shanks 55, Wynn 56, ...] [BD, Padua-Arguelles 23]

A quite simple technique that allows to treat oscillating sums and integrals. Contour deformation not necessary. But reproduces results for integrals treated via contour deformation.

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Works well for sums with actions that are at most linear in the summation variable. Consistent with quantum mechanics (Bohr quantization).





### Action quadratic in discrete variable

Partial sums for an action quadratic in the summation variable



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Partial sums for an action quadratic in the summation variable

True saddle point Pseudo saddle point

Euclidean quantum gravity: rotate conformal factor by hand (if you can)

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Lorentzian path integral: deformation of contour/ Picard-Lefschetz

• Convergence criterium automatically selects 'right' contour

Borissova, BD 23]

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$$\sim \int \exp(+p_{4-1}\lambda^2) \, d\lambda$$
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• Expectation values for (powers of) lengths in spike configurations: -infinite in Euclidean Quantum Gravity

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### Challenge: Which kind of Euclidean simulations/ calculations to trust?

[Mini-superspace: Feldbrugge, Lehners, Turok 17, ... Lorentzian Regge: Asante, BD, Padua-Arguelles 21, Borissova, BD 23]

$$\sim \int \exp(+p_{4-1}\lambda^2) \, d\lambda$$
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What kind of Lorentzian configurations to sum over?

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Lorentzian space time with regular light cone structure.

Lorentzian space time with irregular light cone structure.

![](_page_24_Picture_5.jpeg)

What kind of Lorentzian configurations to sum over?

![](_page_25_Picture_2.jpeg)

Lorentzian space time with regular light cone structure.

Lorentzian space time with irregular light cone structure.

![](_page_25_Picture_5.jpeg)

Valid Euclidean space time.

What kind of Lorentzian configurations to sum over?

![](_page_26_Picture_2.jpeg)

Lorentzian space time with regular light cone structure.

Lorentzian space time with irregular light cone structure.

### Causal Dynamical Triangulations

does forbid such configurations. Leads to a new universality class of random triangulations and an interesting continuum limit in 4 dimensions. [Ambjorn, Loll 1998+] [without preferred slicing: Jordan, Loll 2013]

![](_page_26_Picture_7.jpeg)

Valid Euclidean space time.

What kind of Lorentzian configurations to sum over?

![](_page_27_Picture_2.jpeg)

Lorentzian space time with regular light cone structure.

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![](_page_27_Picture_6.jpeg)

Lorentzian geometry fully encoded in causal relations and volume=number of points.

[from Surya 2019]

Causal sets: Configurations with irregular light cone structure are allowed.

[Talks by Carlip, Surya]

![](_page_27_Figure_11.jpeg)

![](_page_27_Figure_12.jpeg)

What kind of Lorentzian configurations to sum over?

![](_page_28_Picture_2.jpeg)

Lorentzian space time with regular light cone structure. Lorentzian space time with irregular light cone structure.

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Causal sets:

[Talks by Carlip, Surya]

Configurations with irregular light cone structure are allowed.

Imaginary terms in action?

![](_page_28_Figure_14.jpeg)

![](_page_28_Figure_15.jpeg)

![](_page_29_Picture_0.jpeg)

Configuration space: lengths (or other geometric quantities) associated to edges of triangulation.

# Regge gravity (and spin foams)

![](_page_30_Picture_0.jpeg)

Configuration space: lengths (or other geometric quantities) associated to edges of triangulation.

Cr

Allowed Lorentzian triangulations, all edge lengths are space-like.

![](_page_30_Picture_3.jpeg)

![](_page_30_Figure_4.jpeg)

# Regge gravity (and spin foams)

![](_page_30_Figure_6.jpeg)

![](_page_30_Figure_7.jpeg)

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Configuration space: lengths (or other geometric quantities) associated to edges of triangulation.

Allowed Lorentzian triangulations, all edge lengths are space-like.

![](_page_31_Figure_3.jpeg)

![](_page_31_Figure_4.jpeg)

Configurations with irregular light cone structure appear to be generic. Appear even in configurations describing mini-superspace cosmology. Results indicated that such configurations should be included in the path integral.

# Regge gravity (and spin foams)

[Asante, BD, Padua-Arguelles 2022]

![](_page_31_Picture_9.jpeg)

![](_page_32_Picture_0.jpeg)

![](_page_32_Picture_2.jpeg)

[Sorkin 1974, Sorkin 2019] [Jia 2022] [Asante, BD, Padua-Arguelles 2021]

![](_page_32_Figure_4.jpeg)

![](_page_33_Picture_0.jpeg)

![](_page_33_Picture_2.jpeg)

Irregular light cone configurations (of co-dimension 2) lead to branch cuts for the complex Regge action and imaginary contributions with opposite signs. [Asante, BD, Padua-Arguelles 2021]

## Complex Regge action

[Sorkin 1974, Sorkin 2019] [Jia 2022] [Asante, BD, Padua-Arguelles 2021]

![](_page_33_Figure_8.jpeg)

![](_page_34_Picture_2.jpeg)

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[Talk by Asante]

time like Variable

+ SEUCL

enhaucin

## **Complex Regge action**

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![](_page_34_Figure_11.jpeg)

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## **Complex Regge action**

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![](_page_35_Figure_11.jpeg)

Which choice of sign for the path integral?

Suppressing sign: Mechanism for suppressing light cone irregular configurations.

Possibility to reconcile CDT and Causal Sets positions?

![](_page_35_Figure_15.jpeg)

![](_page_35_Figure_16.jpeg)

![](_page_35_Picture_17.jpeg)

Do such imaginary terms appear in the continuum?

![](_page_36_Picture_2.jpeg)

Do such imaginary terms appear in the continuum?

![](_page_37_Picture_2.jpeg)

Complexification needed for discussion of topology change.

Which complexified metrics should be included in the path integral?

[Louko-Sorkin 95]

[Witten 21, Lehners 21, ...]

Do such imaginary terms appear in the continuum?

![](_page_38_Picture_2.jpeg)

[Louko-Sorkin 95] Complexification needed for discussion of topology change. Which complexified metrics should be included in the path integral? [Witten 21, Lehners 21, ...] Complex (boundary) terms in continuum action. [Neiman 13] Gravitational thermodynamics from Lorentzian path integrals: [Marolf 22] Need to include co-dimension 2 singularities. Also for replica copies.

These are already there in simplicial approaches/ Regge gravity!

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Thermodynamics partition functions require contour along enhancing side of branch cut.

[BD, Jacobson, Padua-Arguelles, TA]

![](_page_39_Picture_9.jpeg)

Do such imaginary terms appear in the continuum?

![](_page_40_Picture_2.jpeg)

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Which complexified metrics s

Complex (boundary) terms

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Thermodynamics partition functions require contour along enhancing side of branch cut.

Challenges: Understand better role of complex structures for Lorentzian gravitational path integral. **Canonical Formalism?** Should we include configurations describing topology change? Complex actions for sum over topologies?

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should be included in the path integral?	[Witten 21, Lehners 21, .
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[BD, Jacobson, Padua-Arguelles, TA]

![](_page_40_Picture_11.jpeg)

## Lorentzian quantum gravity

- More techniques for Lorentzian path integrals
- Lorentzian configurations have additional light cone structure: can be (often) irregular
- Irregular light cone structure leads to branch cuts for the gravitational action and imaginary terms
- Which Lorentzian configurations to include in the path integral? Which side of the branch cut?
- Thermodynamical interpretation?

### What is the structure of quantum space time?

Quantum Field Theory:

Fields  $\phi(x, t)$  are observables. n-point functions allow to reconstruct full theory.

How much can we resolve space-time?

![](_page_43_Picture_5.jpeg)

Quantum Gravity:

Fields  $\phi(x, t)$  are not observables. No n-point functions.

How much can we resolve space-time?

![](_page_44_Picture_5.jpeg)

Quantum Gravity:

Relational observables [Einstein, ...]

Fields  $\phi(x, t)$  are not observables. No n-point functions.

 $\phi(\psi_1,\psi_2,\psi_3,\psi_4)$ 

How much can we resolve space-time?

[Talks by Kiefer, Ferrero, Giacomini, Hoehn, Menendez-Pidal, Rastgoo, Ruf, ...]

![](_page_45_Picture_8.jpeg)

Quantum Gravity:

Relational observables [Einstein, ...]

 $\phi(\psi_1,\psi_2,\psi_3,\psi_4)$ 

But there seem to be no perfect clocks in our (relativistic) universe.

![](_page_46_Picture_7.jpeg)

How much can we resolve space-time?

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[Talks by Kiefer, Ferrero, Giacomini, Hoehn, Menendez-Pidal, Rastgoo, Ruf, ...]

[Giddings, Marolf, Hartle 05; BD, Tambornino 06, Giddings, Donnely 15; Hoehn et al 20+] Chaos and other aspects: [Bojowald, Hoehn, et al. 12+, BD, Hoehn, Kosolowski, Nelson 17] End of time: [Talks by Gielen, Menendez-Pidal]

![](_page_46_Picture_12.jpeg)

![](_page_46_Figure_15.jpeg)

### $\left[\phi(\psi), \phi(\psi + \epsilon)\right] = G(\psi, \psi + \epsilon) \left(1 + \frac{\operatorname{Energy}(\phi)}{\operatorname{Energy}(\psi)}\right)$ Bound on resolution with relativistic clocks:

[Giddings, Marolf, Hartle 05, BD, Tambornino 06]

How much can we resolve space-time?

![](_page_47_Picture_5.jpeg)

Increasing energy (and size) of clock field leads to black holes.  $\Rightarrow$ Locality Bounds

![](_page_47_Picture_7.jpeg)

![](_page_47_Picture_8.jpeg)

![](_page_47_Picture_9.jpeg)

Bound on resolution  
with relativistic clocks: 
$$\left[\phi(\psi), \phi(\psi + \epsilon)\right] = G(\psi, \psi + \epsilon) \left(1 + \frac{\text{Energy}}{\text{Energy}}\right)$$

[Giddings, Marolf, Hartle 05, BD, Tambornino 06]

A new uncertainty principle?

New uncertainty for Time of arrival operator $\Delta t > 1/\overline{E}$ .	[4
The same holds for relativistic clocks.	[

How much can we resolve space-time?

![](_page_48_Picture_7.jpeg)

Increasing energy (and size) of clock field leads to black holes.  $\Rightarrow$ Locality Bounds

![](_page_48_Picture_9.jpeg)

[Aharanov et al 97]

BD, Tambornino 06]

![](_page_48_Picture_12.jpeg)

![](_page_48_Picture_13.jpeg)

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The same holds for relativistic clocks.

High energy scattering leading to black holes: "The end of short distance physics". UV-IR mixing.

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Aharanov et al 97

[BD, Tambornino 06]

[Banks,Fischler 99; Giddings-Thomas 01, Dvali et al ...]

![](_page_49_Picture_15.jpeg)

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High energy scattering leading to black holes: "The end of short distance physics". UV-IR mixing.

Challenge: Better understanding of type and algebra of observables.

How much can we resolve space-time?

![](_page_50_Picture_9.jpeg)

Increasing energy (and size) of clock field leads to black holes.  $\Rightarrow$ Locality Bounds

![](_page_50_Picture_11.jpeg)

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![](_page_50_Picture_14.jpeg)

![](_page_50_Picture_15.jpeg)

## Structure of quantum space-time

Quantum field theory: space-time is an index set.

Quantum gravity: this index set is dynamical and quantum.

[Talks by Freidel, Giacomini, Hoehn, ... ]

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What is the resulting structure of quantum space time?

- fractal dimensions/ space-times
- non-commutative space-time
- geometric operators (with discrete spectra)
- holography
- local holography
- matrix and tensor models
- group field theories
- strings
- area metrics
- twistor space
- relative locality and Born duality
- •

[Talks by Freidel, Giacomini, Hoehn, ...]

Space time emergent

![](_page_52_Picture_19.jpeg)

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[Talks by Freidel, Giacomini, Hoehn, ...]

Space time emergent

![](_page_53_Picture_20.jpeg)

### Challenge: Observable algebra and symmetries $\leftrightarrow$ Structure of quantum space-time

## Reconstructing our universe

![](_page_54_Picture_1.jpeg)

![](_page_54_Picture_2.jpeg)

Challenge: Bridge enormous number of scales. Understand renormalization and develop (common) language for 'emergence' of space-time.

![](_page_54_Picture_4.jpeg)

## Understanding our universe

"So far we have not seen any signatures of quantum gravity."

"Are we now really at the dawn of quantum gravity phenomenology?"

[Talk by Amelino-Camelia]

## Understanding our universe

"So far we have not seen any signatures of quantum gravity."

"Are we now really at the dawn of quantum gravity phenomenology?"

But there are many features of our universe asking for explanation:

- Quantum theory?
- (Macroscopic) four-dimensional smooth space-time
- Lorentzian signature
- Gravitational dynamics
- Small and positive cosmological constant

[Talk:Yasaman]

[Talk by Amelino-Camelia]

• Inflation?

- (Simple) initial conditions
- Matter
- Dark Matter
- Types of matter and values of matter couplings

[Talk:Wetterich]

![](_page_57_Picture_0.jpeg)