

# Everpresent $\Lambda$ Cosmology

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Work with Santanu Das & Arad Nasiri ([arXiv:2304.03819](https://arxiv.org/abs/2304.03819))

**Quantum Gravity 2023**

**Radboud University, Nijmegen**

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**Imperial College  
London**

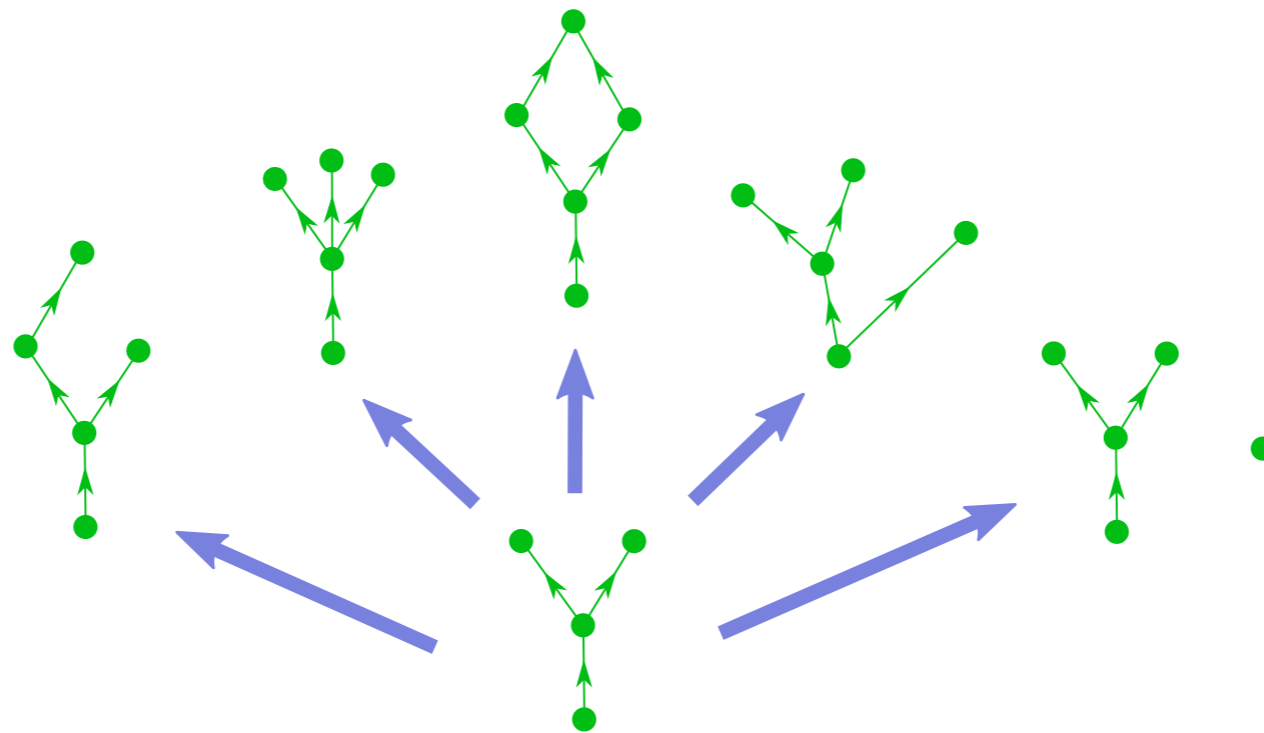
July 12, 2023

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# Causal Set Theory: Growing by Number

In classical dynamical models, a causal set **grows element by element** sequentially. At stage 1, the first element is born, and subsequently at each stage  $n$  we have an  $n$ -element causal set.



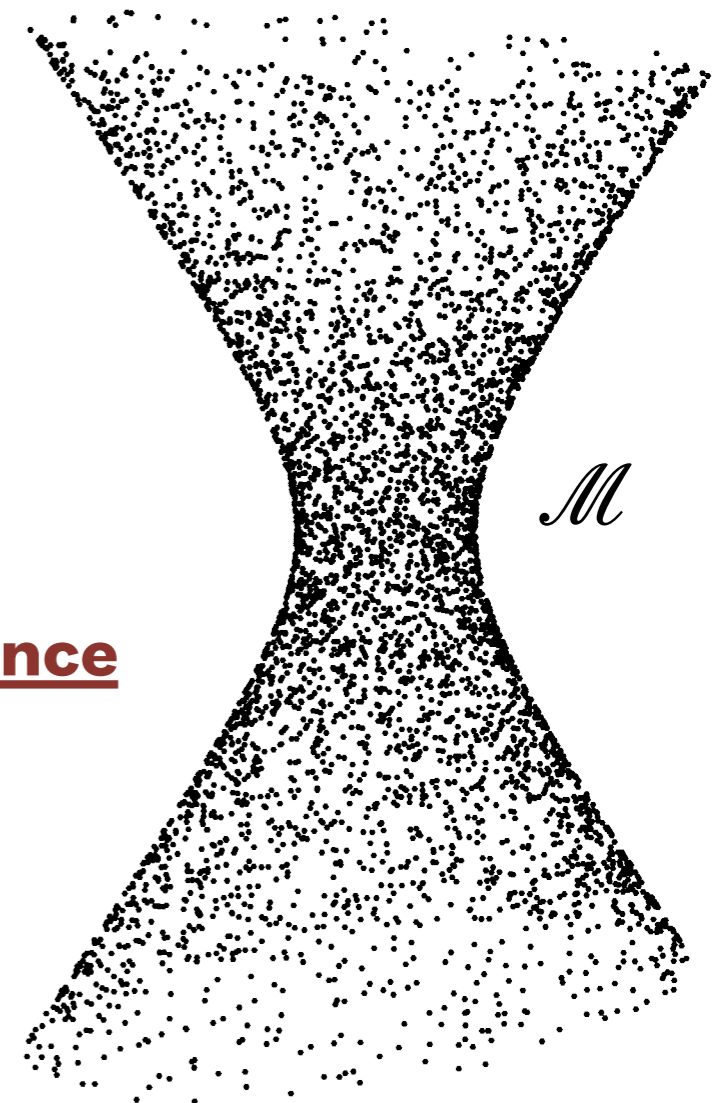
**Dynamics ingredient: number of elements plays a similar role to time**

# Causal Set Theory: Fundamental Discreteness

Causal sets that are approximated by continuum manifolds have a **number-volume correspondence** such that: the number of elements  $N$  within any arbitrary region with volume  $V$  is statistically proportional to  $V$ . The **Poisson distribution** ensures this correspondence with minimal variance.

$$\langle N \rangle = V$$
$$\delta N = \sqrt{V}$$

**Kinematics ingredient: number-volume correspondence according to the Poisson distribution**



# Everpresent $\Lambda$

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**Dynamics input: number of elements plays a similar role to time**

$$\mathcal{Z}(V) \sim \int_{\text{Vol}(\mathcal{M})=V} \mathcal{D}g_{\mu\nu} e^{iS_G[g]}$$

# Everpresent $\Lambda$

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$$\mathcal{Z}(V) \sim \int_{\text{Vol}(\mathcal{M})=V} \mathcal{D}g_{\mu\nu} e^{iS_G[g]} \xrightarrow{\text{Fourier Transform}} \mathcal{Z}(\Lambda) = \int dV e^{-i\Lambda V} \mathcal{Z}(V)$$

$$\mathcal{Z}(\Lambda) \sim \int \mathcal{D}g_{\mu\nu} \exp \left( iS_G[g] - i\Lambda \int d^4x \sqrt{-g} \right)$$

Therefore, spacetime volume  $V$  and  $\Lambda$  are quantum mechanically conjugate

$$\frac{\delta\Lambda}{8\pi G} \cdot \delta V \geq \frac{\hbar}{2}$$

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A causal set with fixed  $N$  can be approximated by continuum spacetimes with mean  $\langle V \rangle = N$  and standard deviation  $\delta V = \sqrt{N} \sim \sqrt{V}$ . Also:  $N, V \gg 1$

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$$\delta\Lambda \delta V \sim 1 \quad \xrightarrow{\text{Assume } \langle \Lambda \rangle = 0} \quad \Lambda \sim \delta\Lambda \sim 1/\delta V = 1/\sqrt{V} \sim H^2 \sim 10^{-121}!$$

# A Phenomenological Model of Everpresent $\Lambda$

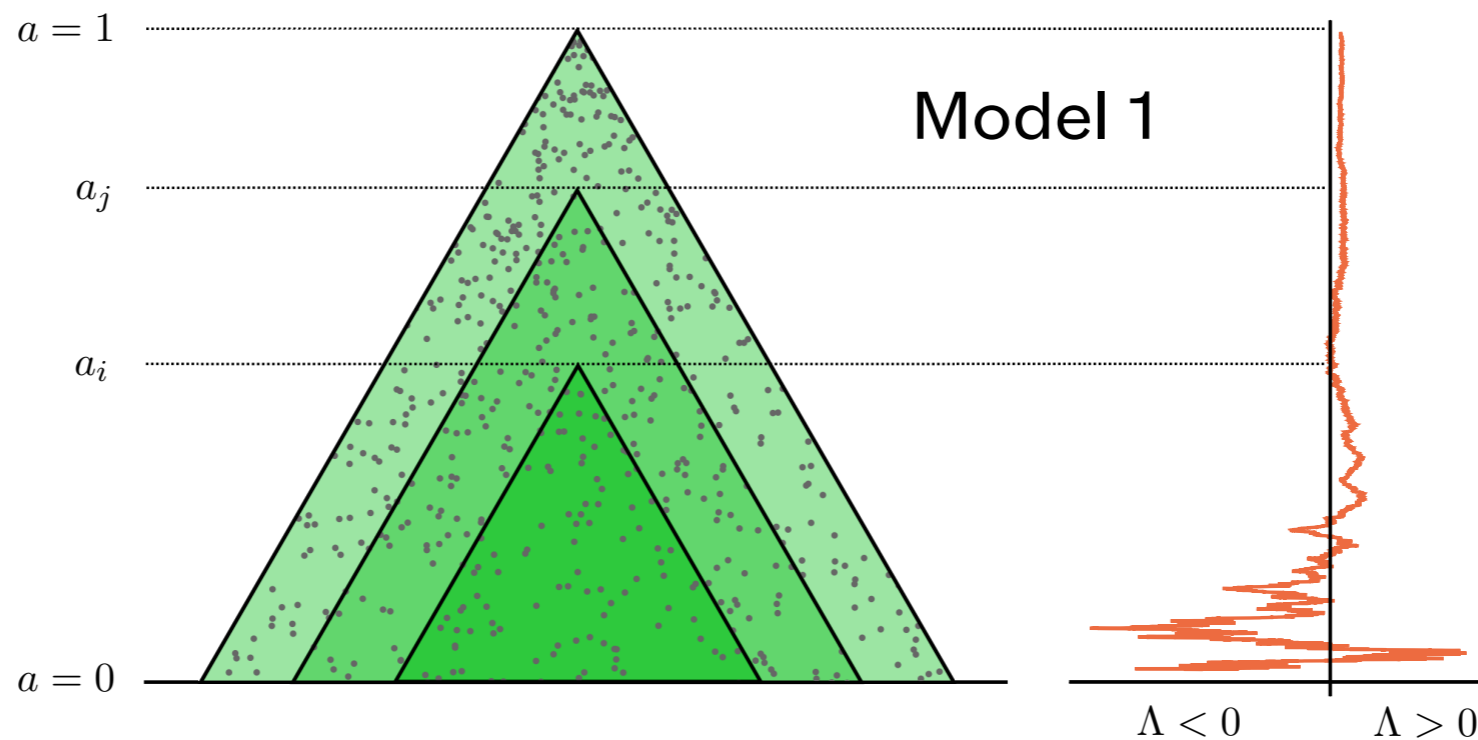
Ahmed, Dodelson, Greene, and Sorkin, "Everpresent  $\Lambda$ ", PRD 69, no. 10 103523, 2004.

$$|\Lambda| \sim 1/\sqrt{V}$$

Each causal set element contributes a random variable with standard

deviation  $\alpha$  to  $S_\Lambda$ :

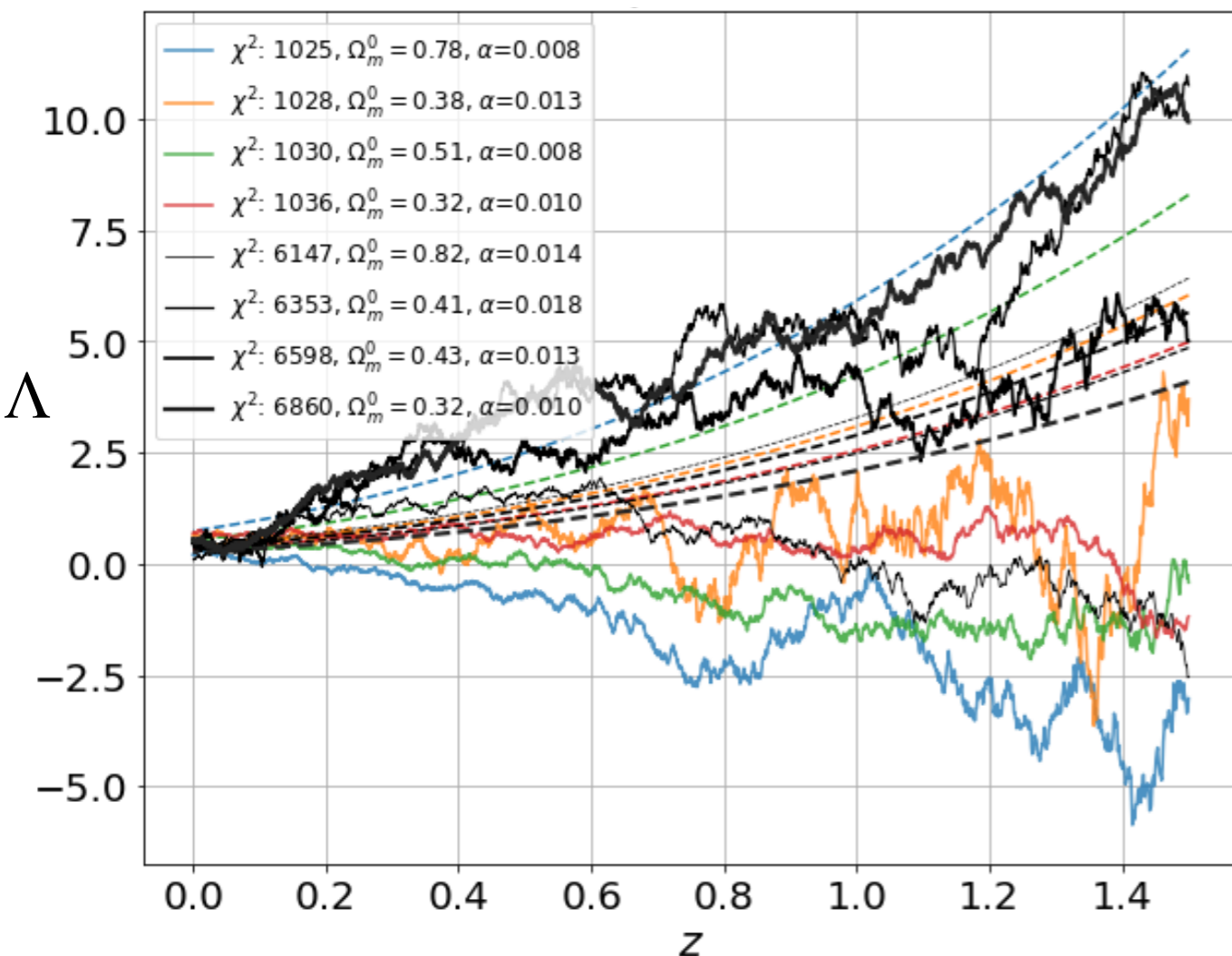
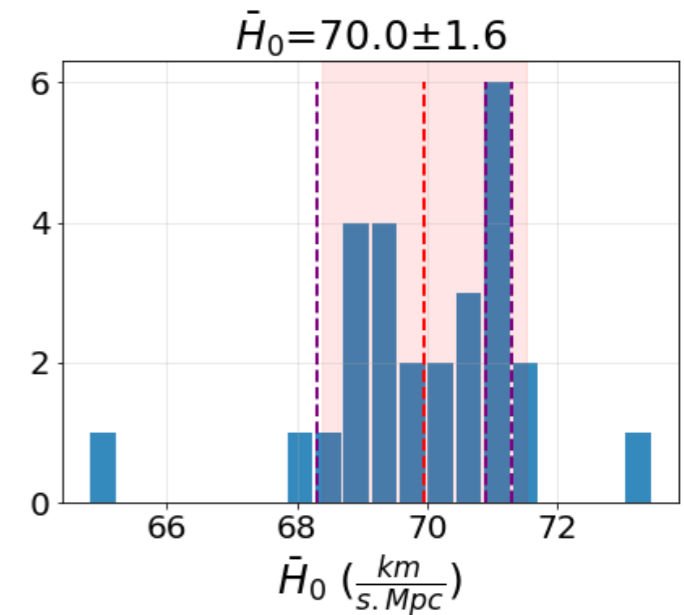
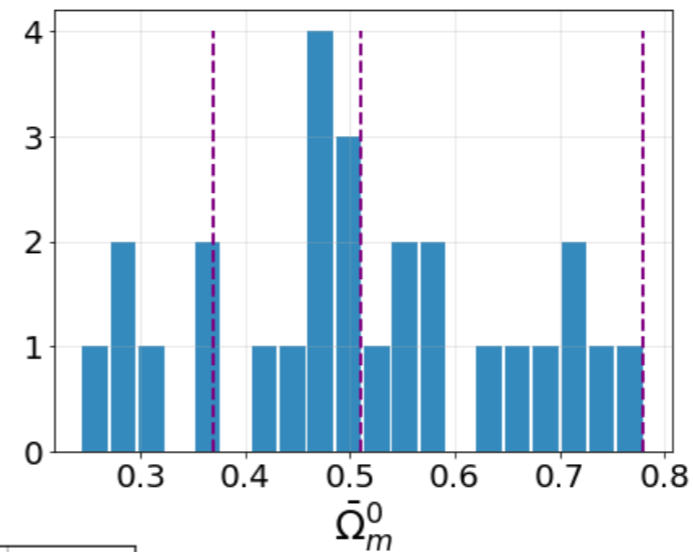
$$S_\Lambda |_{1 \text{ element}} = \begin{cases} \alpha, & p = \frac{1}{2} \\ -\alpha, & p = \frac{1}{2} \end{cases}$$





# Type Ia Supernovae

Out of a sample of 20,000 seeds,  
3 Everpresent  $\Lambda$  histories  
(dashed lines) yielded a better  $\chi^2$   
(1025, 1028, 1030) than  $\Lambda$ CDM (1033).



Dark energy densities that are smaller than matter density at small redshifts (colored) are favored over those with comparable or larger dark energy densities (black).

# Summary & Outlook

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- Everpresent  $\Lambda$ : a fluctuating cosmological constant from spacetime discreteness
- Many open questions remain: e.g. why is the mean 0?
- Improvements: CMB data fit, quantum modifications, stochastic differential equations, incorporation of inhomogeneities