



University
of Bremen



QuantumFrontiers
Cluster of Excellence



CENTER OF
APPLIED SPACE TECHNOLOGY
AND MICROGRAVITY



GENERAL RELATIVITY MEETS GEODESY

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for the TerraQ collaboration

The 5th EPS Conference on Gravitation
Unlocking Gravity through Computations
Prague, 09-11 December 2024



Irrigation

Droughts



Flooding

Ice melting

Sea level change

Irrigation

Droughts

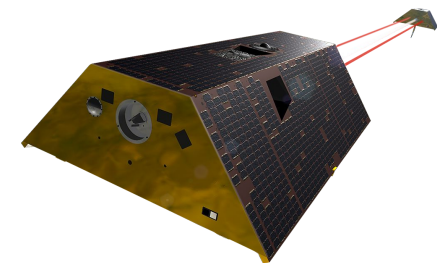
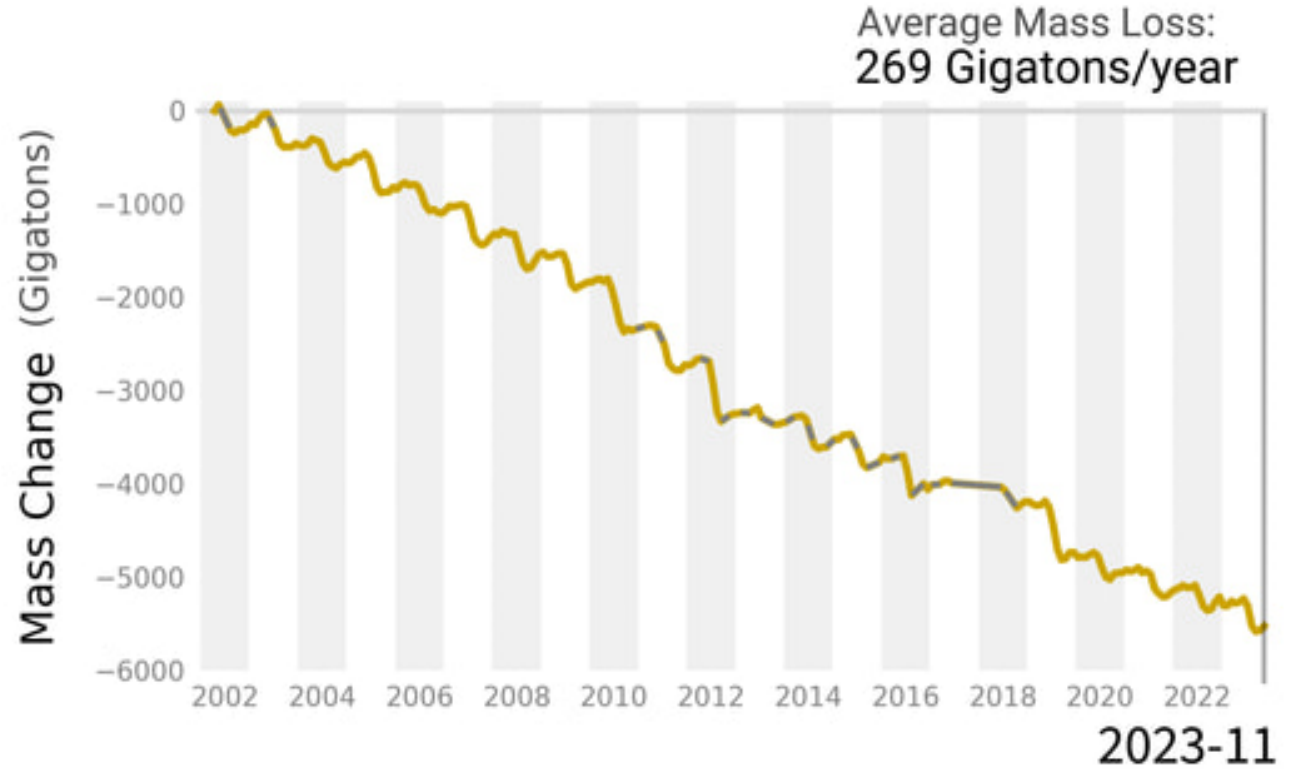
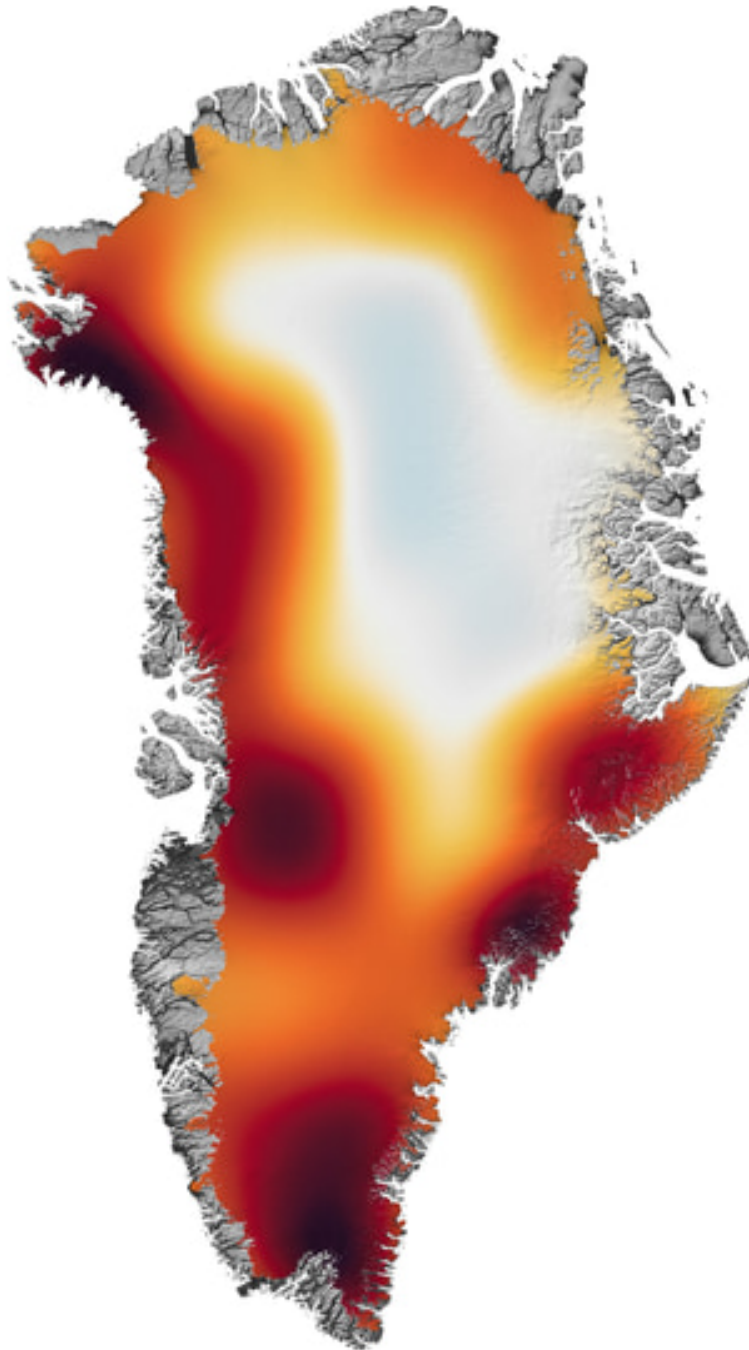


Flooding

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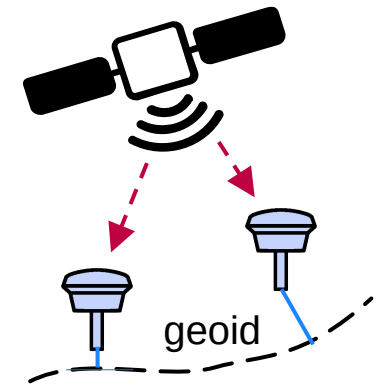
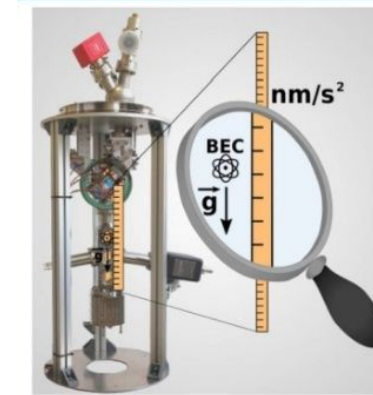
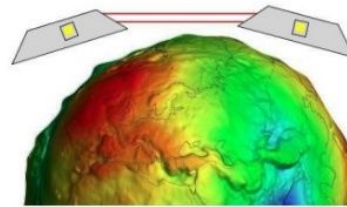
GRACE AND GRACE-FO Observations of Greenland Land Ice Mass Changes



How to observe the gravity field?

Available techniques

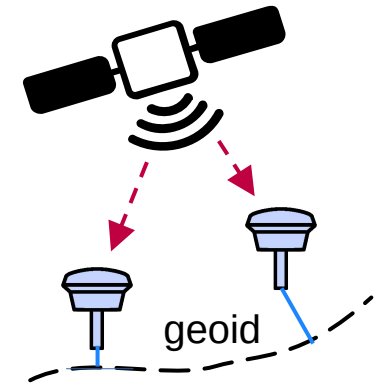
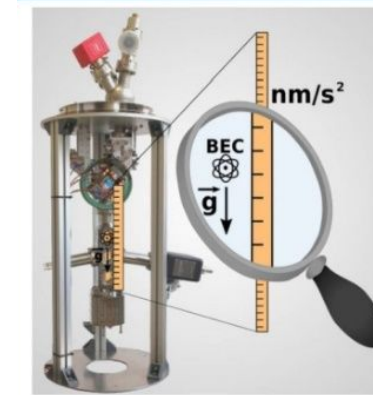
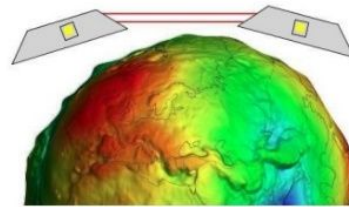
- Satellite gravimetry
- Terrestrial gravimetry
- GNSS station displacements



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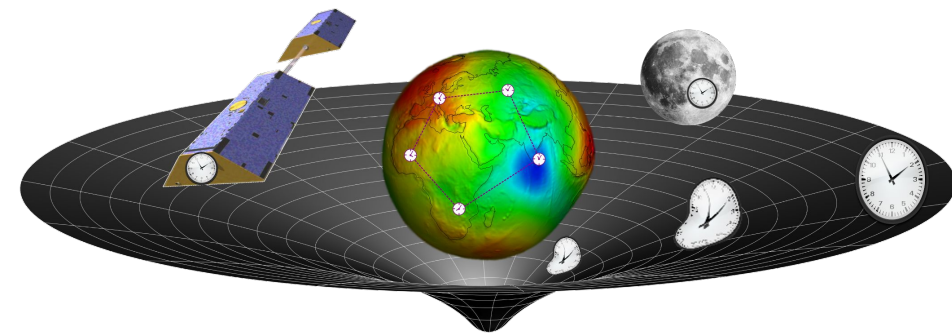
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General Relativity

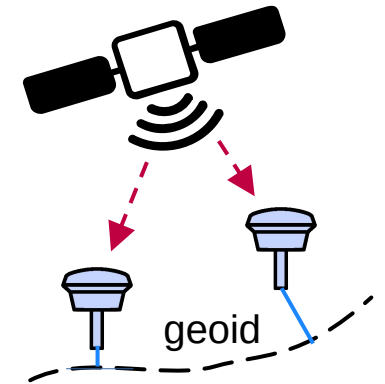
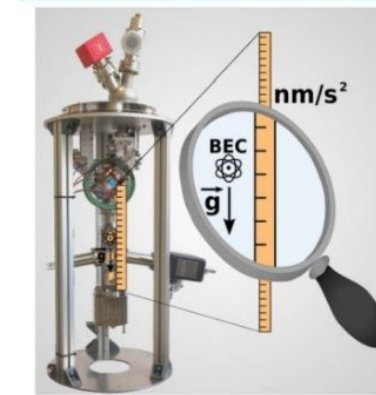
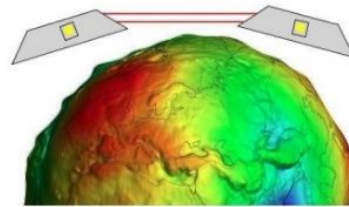
- Sets the stage for all measurements
- Completely new measurement approach:
 - clocks probe the curved spacetime geometry by their proper time



How to observe the gravity field?

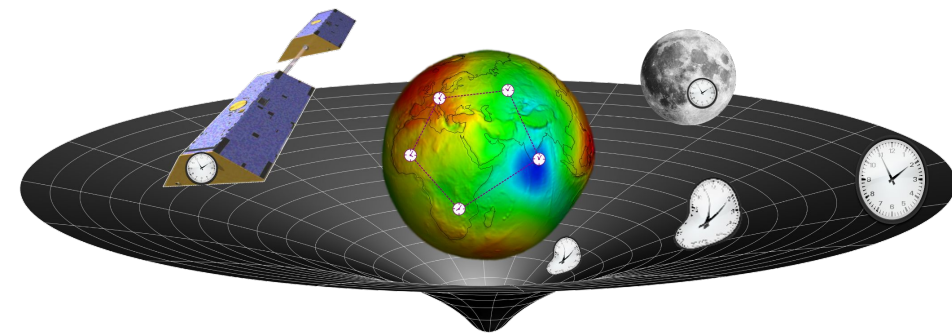
Available techniques

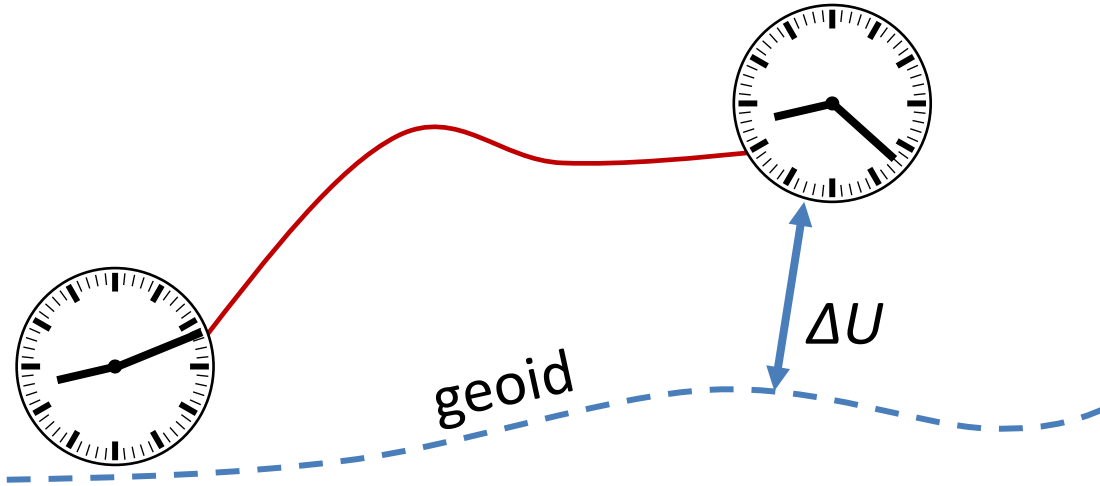
- Satellite gravimetry
- Terrestrial gravimetry
- GNSS station displacements
- **Clock networks**



General Relativity

- Sets the stage for all measurements
- Completely new measurement approach:
 - clocks probe the curved spacetime geometry by their proper time





gravitational red shift:

$$\Delta\nu/\nu = \Delta U/c^2$$

M. Vermeer, Rep. of the Finnish Geod. Insti. (1983)
A. Bjerhammar, Bull. Geodesique (1985)

Clock comparison

- 1 cm-resolution within reach (1×10^{-18} optical clock)
- no error accumulation over distance
- high spatial resolution (atoms are small)
- fast measurements

Requirements

- two (transportable) optical clocks with 10^{-18} uncertainty
- link to compare them
- physical justification → GR

Chronometric levelling – state of the art

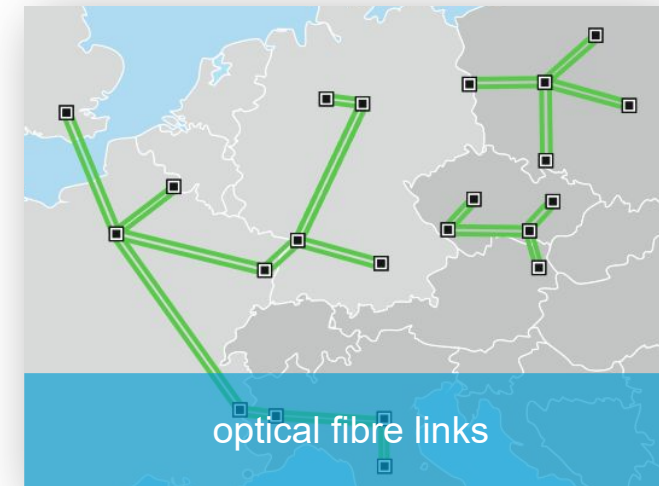
Transportable optical clocks

- 2nd generation Sr lattice clock
 - most stable transportable clock laser
[Herbers *et al.* Opt. Lett. \(2022\)](#)
 - uncertainty $<3 \times 10^{-18}$
[Lisdat *et al.* Phys. Rev. Res. \(2021\)](#),
[Dörscher *et al.* Phys. Rev. Res. \(2023\)](#)



Frequency links

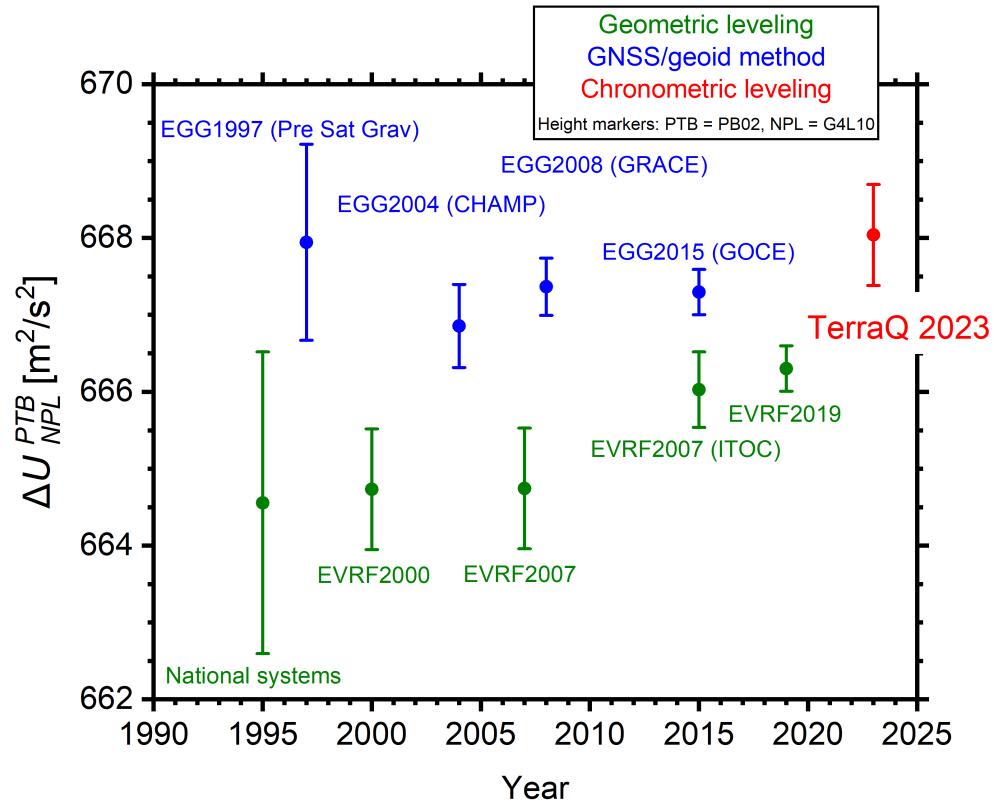
- Optical fibre links:
 - $<10^{-18}$ instability in 100 s,
 $<10^{-19}$ offsets
[Schioppo *et al.* Nature Commun \(2022\)](#),
[Koke *et al.* New J. Phys. \(2019\)](#)
- GNSS frequency transfer
 - 5×10^{-17} instability on 50 km baseline
[Proc. 55th APTTISA Meeting 2024](#)



Chronometric levelling campaigns

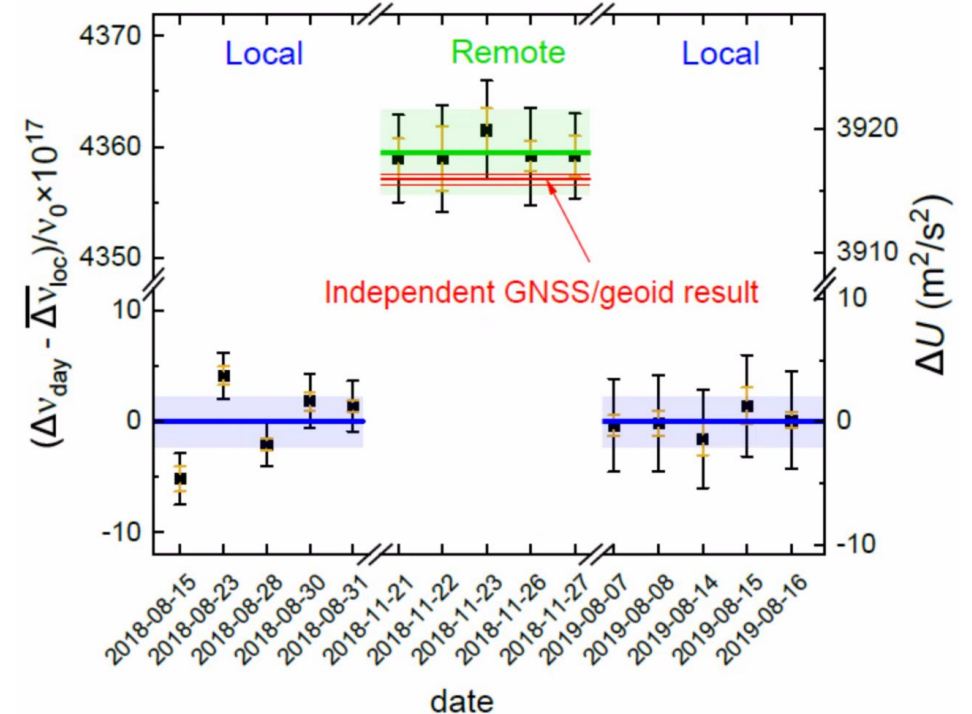
NPL - PTB

- Transportable & laboratory Sr clocks



Munich - PTB

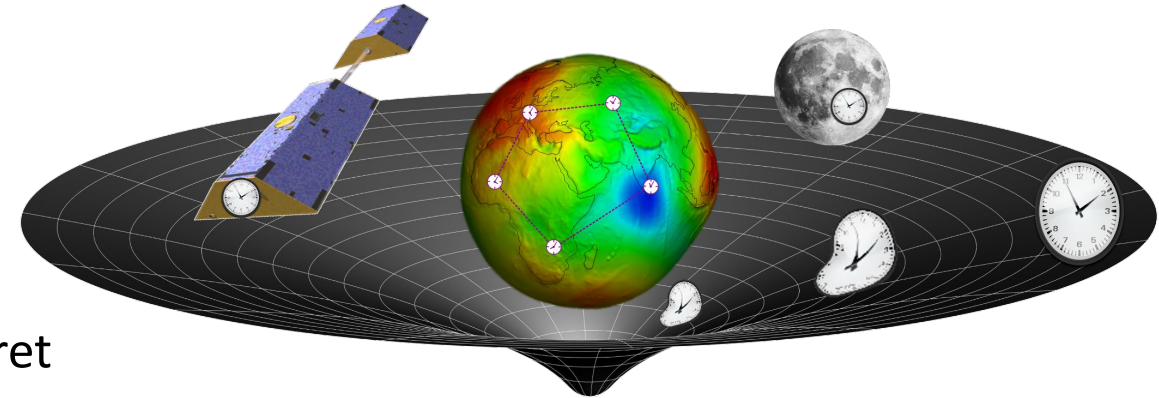
- 2018: uncertainty of 27 cm
Dörscher et al. Phys. Rev. Res. (2023), Lisdat et al. Phys. Rev. Res. (2021)
- 2024: 5 cm goal



General Relativity in Geodesy

- Geodesy is based on (post-)Newtonian notions
 - Reference systems and surfaces
 - Height definitions
 - Newtonian gravity potential only
- ➔ Develop consistent GR theoretical framework to interpret the observations

- Relativity in terrestrial and satellite gravimetry
 - Terrestrial clock networks for realizing a global height system
 - Clocks for gravity field recovery (Earth, space, hybrid)
 - Needs also relativistic effects on satellite motion and on (quantum) sensor



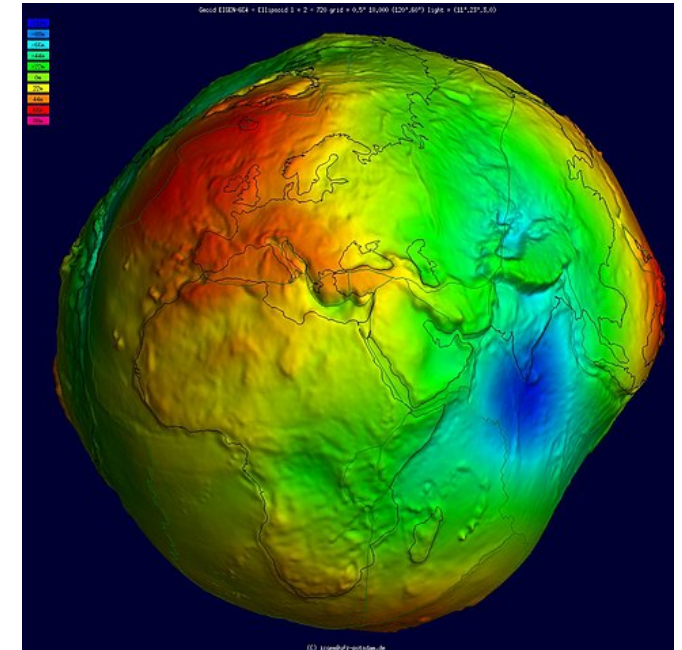
GR framework of Geodesy

Reference surfaces

- Geodesy uses a number of reference surfaces
- One of them is the geoid: the „mathematical figure of Earth“ (Gauss)
 - Equipotential surfaces of the Newtonian gravity potential $W = U + V$
 - One of them („mean sea level“) is singled out by convention $W = W_0$
- Based on measurement of acceleration (a-geoid)

Relativistic geoid

- “Surface where precise clocks run with the same speed“
[A. Bjerhammar, Bull. Geodesique \(1985\)](#)
→ Surface where atomic clocks show vanishing redshift
- Based on measurement of time/frequency (u-geoid)
[M. Soffel et al, Manuscripta Geodaetica \(1988\)](#), [Kopeikin et al, Phys. Lett. A \(2015\)](#)



GR framework of Geodesy

The relativistic geoid Philipp et al, Phys. Rev. D (2017)

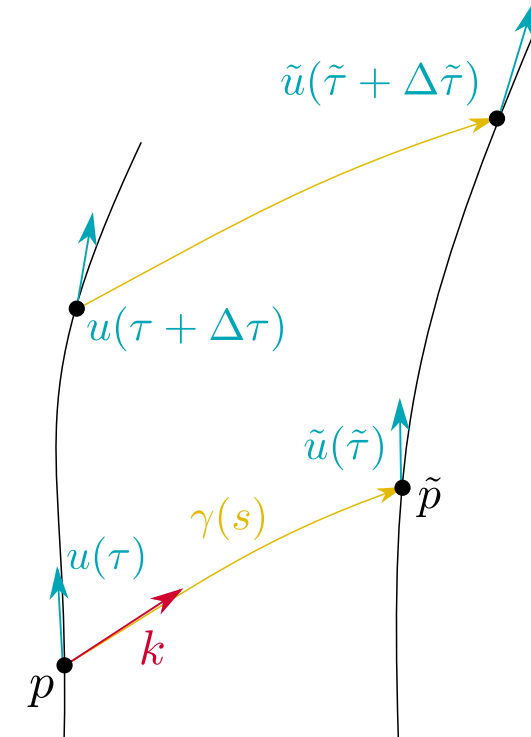
- Introduce congruence of timelike worldlines with four-velocity u
- The redshift z is

$$z + 1 = \frac{d\tilde{\tau}}{d\tau} = \frac{g_{\mu\nu} k^\mu u^\nu|_p}{g_{\rho\sigma} k^\rho u^\sigma|_{\tilde{p}}}$$

- Defines (dimensionless) redshift potential ϕ as $z + 1 = \exp \Delta\phi$
- ϕ is time independent iff $e^\phi u = \xi$ is a Killing vector field Hasse & Perlick, J. Math Phys (1988)
- Equipotential surfaces of ϕ are isochronometric: clocks have vanishing redshift

Equivalence to a-geoid

- We assume rigid rotation, constant angular velocity, no external forces
- allows to introduce an acceleration potential χ Ehlers 1961, Salzmann & Taub, Phys. Rev. (1954)
- We can show: $\phi = \chi \rightarrow$ data fusion applies!



GR framework of Geodesy

Relativistic gravity potential Philipp et al, Phys. Rev. D (2020)

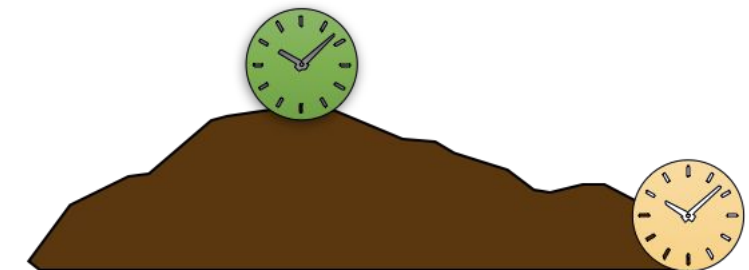
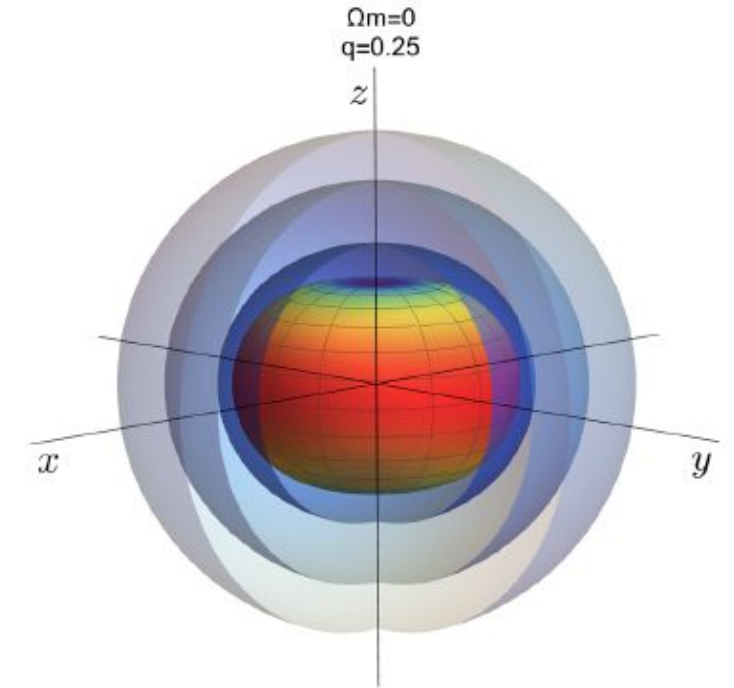
- In adapted coordinates: $\exp(2\phi) = g_{00}$
- Introduce relativistic gravity potential $U^* = c^2(\sqrt{-g_{00}} - 1)$
- The relativistic geoid is then a level surface of U^*

- In ppN limit
$$U^* \approx W + \frac{U(\beta - 1/2)}{c^2}$$

- The redshift
$$z + 1 = \frac{1 + U_2^*/c^2}{1 + U_1^*/c^2} \approx \frac{\Delta W}{c^2}$$

Based on gravity potential

- The (relativistic) **potential numbers** $C_P^* := U_P^* - U_0^* = z(c^2 + U_0^*)$
- The (relativistic) **chronometric height** $H_P^* := C_P^*/\bar{a}$



Comparison relativistic vs Newtonian geoid

Setup

- We choose a simple quadrupolar model of the Earth
- Compare Newtonian geoid based on W with 1st-order post-Newtonian geoid based on U^*
- We expect
 - an overall spherical relativistic contribution due to the monopole
 - a latitude dependent correction (about three orders of magnitude smaller)

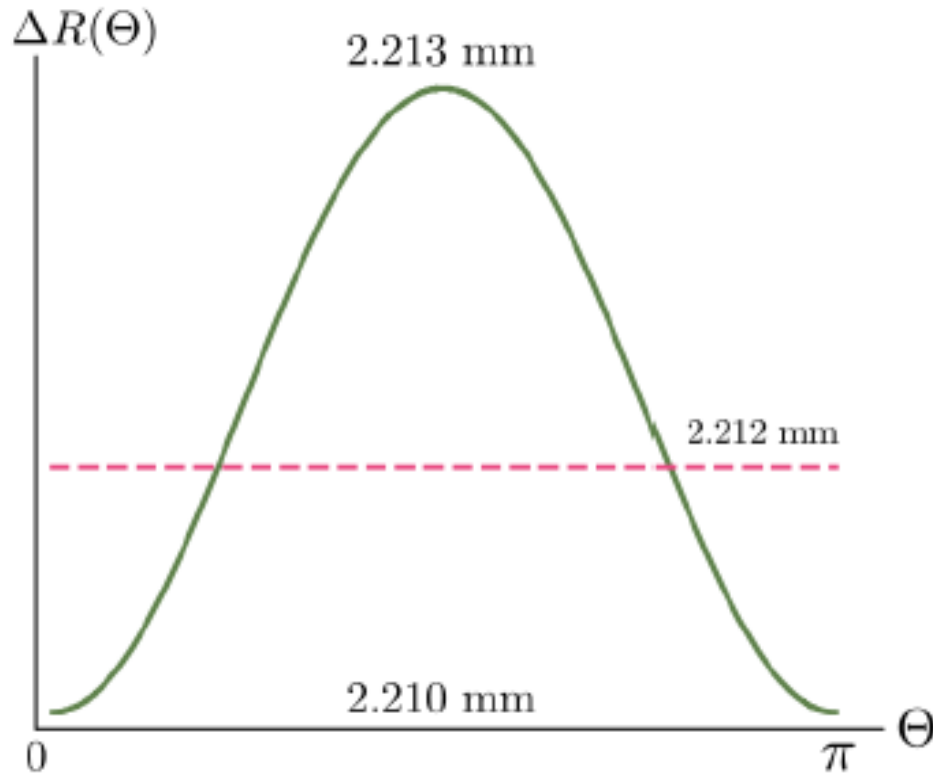
Comparison

- Coordinates cannot be directly compared
→ isometric embedding into Euclidian space
- How to choose U_0^* ?

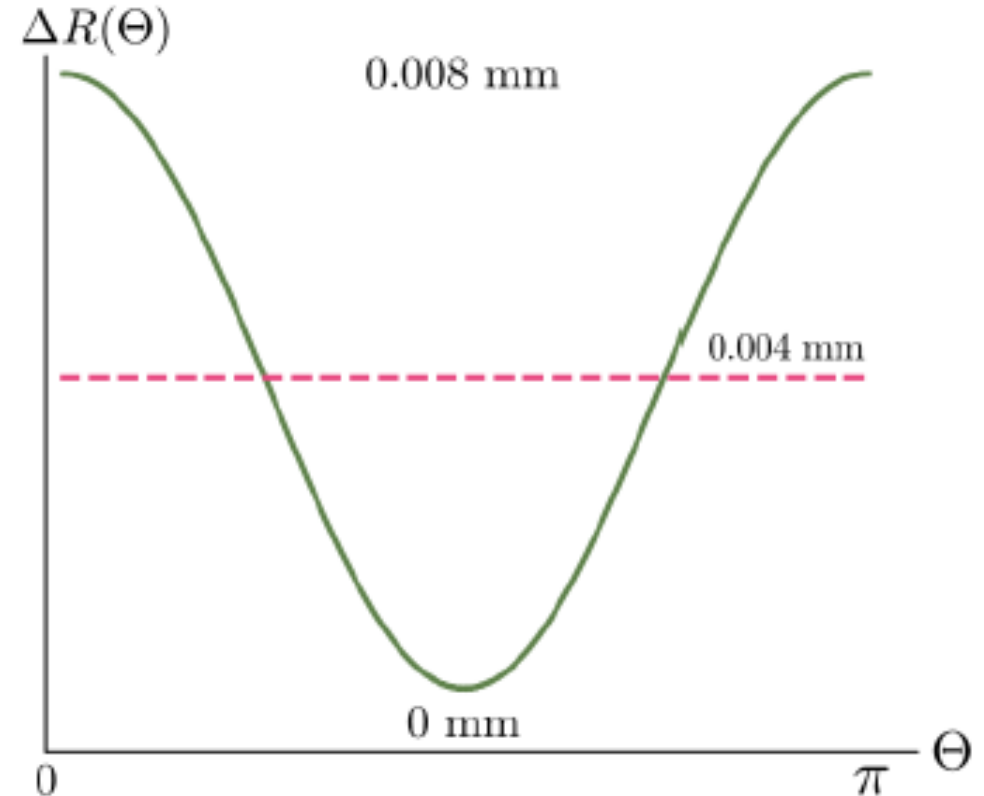
Remember: $U^* \approx W + U/(2c^2)$

$$U_0^* = W_0$$

U_0^* such that geoids coincide at equator



Mean difference: 2.2mm
 Latitudinal variations: $3\mu\text{m}$



Mean difference: $4\mu\text{m}$
 Latitudinal variations: $8\mu\text{m}$

Non-Newtonian gravitational degrees of freedom

Gravitomagnetic effects

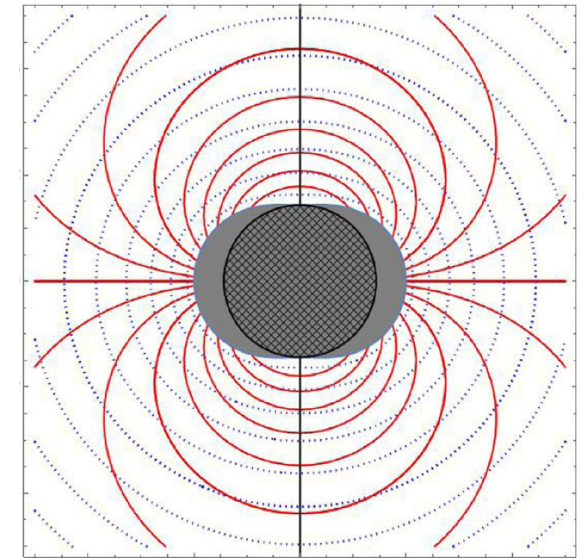
- Rotation also contributes to the gravitational field
→ frame dragging/gravitomagnetism

$$g = g_{00}c^2 dt^2 + 2g_{0i}c dt dx^i + g_{ij} dx^i dx^j$$

- Proceed similar to redshift potential
 - Introduce congruence of timelike Killing observers ξ
 - Then the twist vector field is $\omega^\mu = \epsilon^{\mu\nu\rho\sigma} \xi_\nu \partial_\rho \xi_\sigma$
 - Einstein's vacuum equations imply the twist potential

$$\omega_\mu = \partial_\mu \omega$$

- One equipotential surface may be used as „rotoid“
- Redshift potential ϕ and twist potential ω
→ purely chronometric reference frame

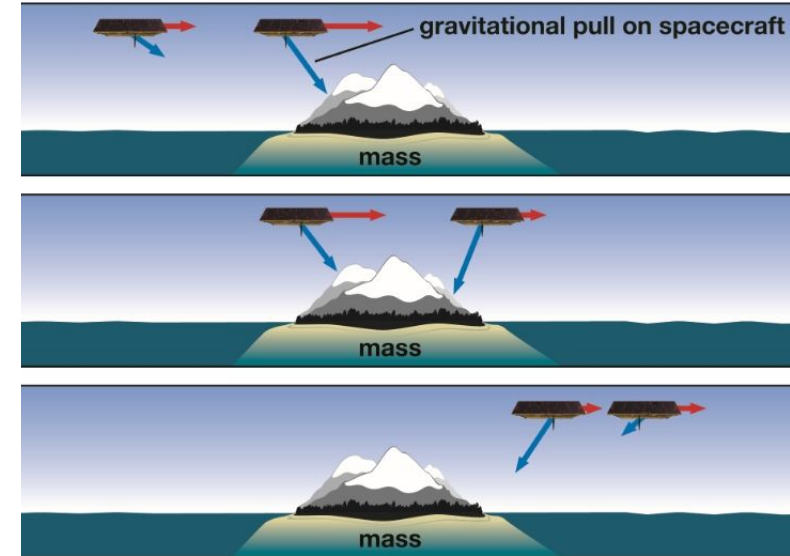


Lämmerzahl & Perlick, Phys. Rev. D (2023)

Satellite gravimetry: GRACE-FO

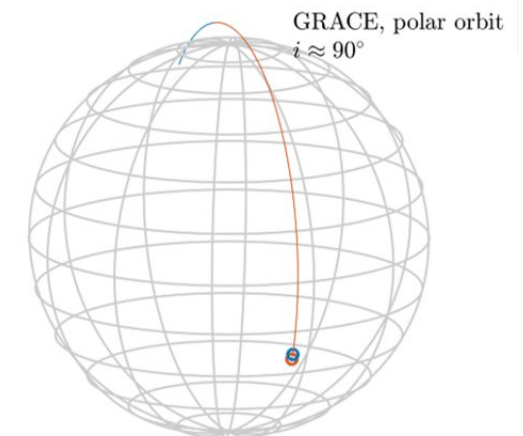


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The mission

- 2 satellites about 220 km apart; Microwave and Laser link
- Range changes are measured on the nm level
- Monthly global gravity field solutions of ~300 km spatial resolution



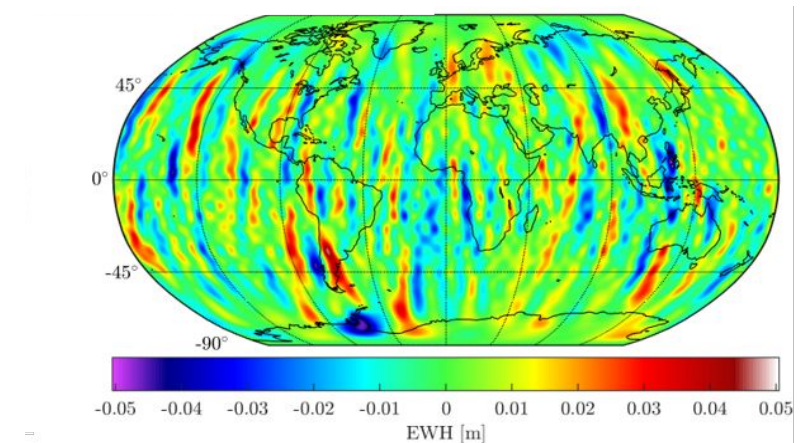
Gravity Field Recovery from space

Basic approach

- The Newtonian gravitational potential U is written as

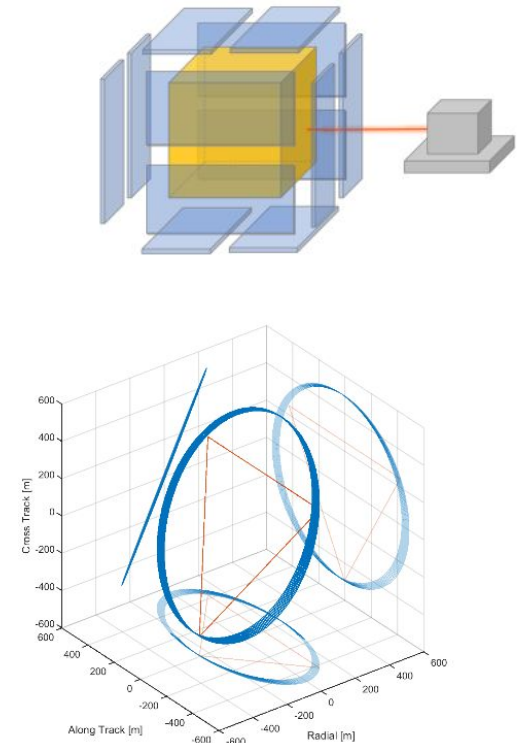
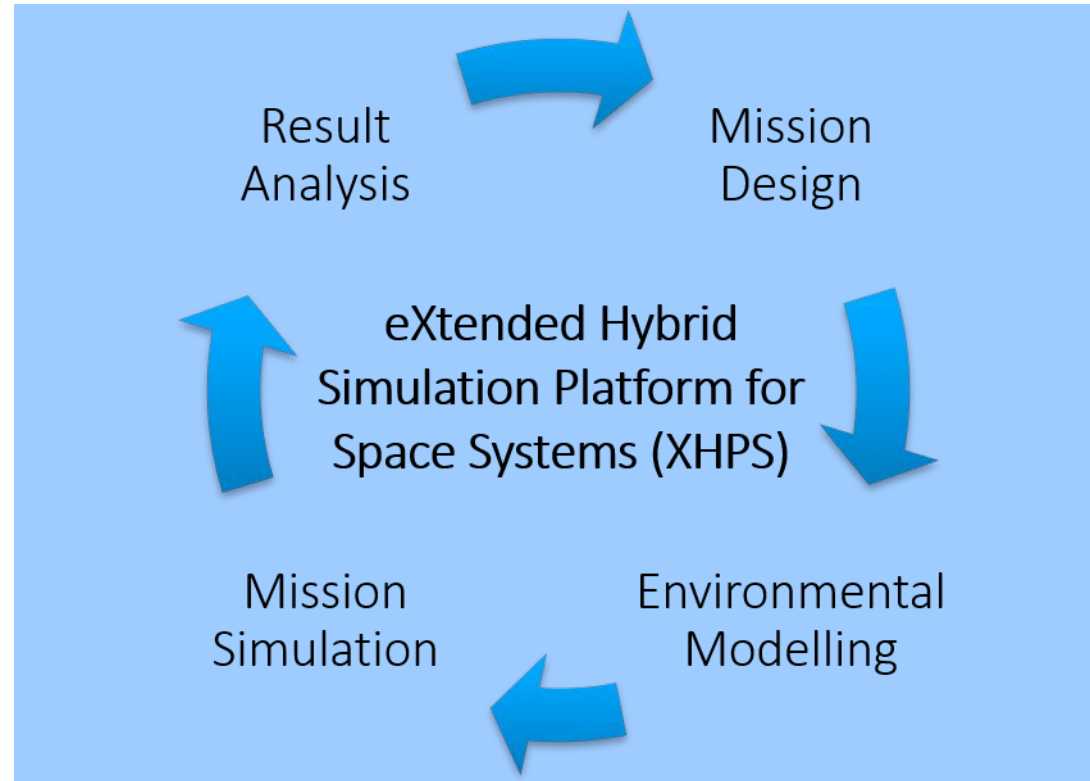
$$U = \frac{GM}{r} \sum_{l=0}^{\infty} \sum_{m=0}^l \left(\frac{R_{\text{ref}}}{r} \right)^l P_{lm}(\sin \theta) \{C_{lm} \cos(m\varphi) + S_{lm} \sin(m\varphi)\}$$

- Goal: determine the coefficients C_{lm} , S_{lm} from the data
 - Model the satellite trajectories
 - Model all kinds of perturbations
 - Model the sensors



Simulator Tool XHPS

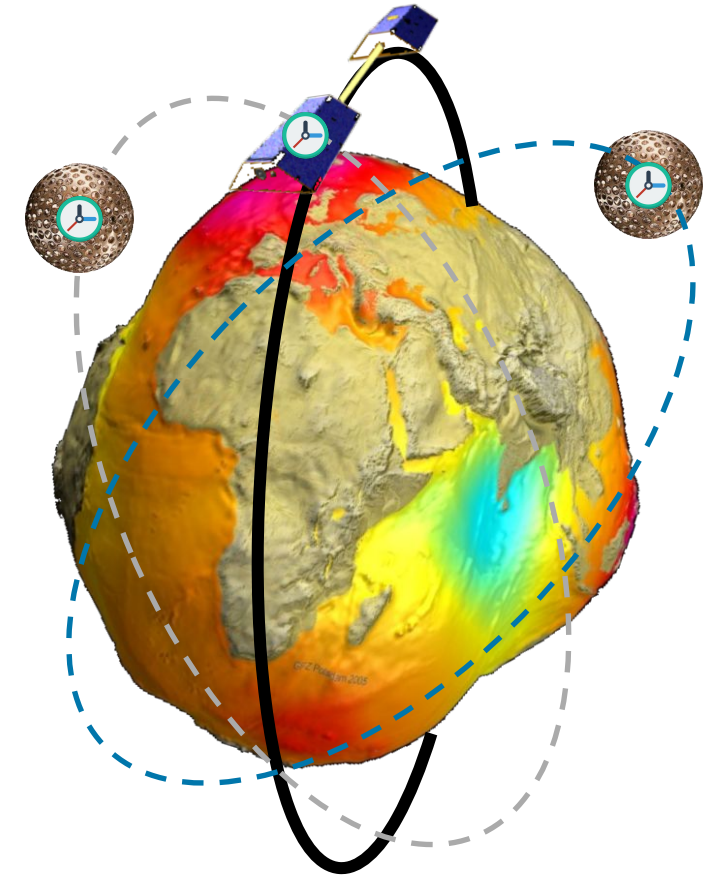
- model of optical-electrostatic accelerometers
- satellite swarm scenarios
- test mass dynamics and a generic drag-free approach
- 1st post-Newtonian order



Clocks in space

Can clocks be used for Gravity Field Recovery from space?

- Challenge: separate the much larger special relativistic effects (satellites move with ~ 10 km/s)
- 1st order Doppler effect can be eliminated by using a two way link
- Uncertainty in 2nd order Doppler effect maps into gravitational potential determination
- Rough estimate:
1 cm geoid corresponds to about $10 \mu\text{m/s}$ in LEO (Low Earth Orbit)
- This seems to be very challenging...



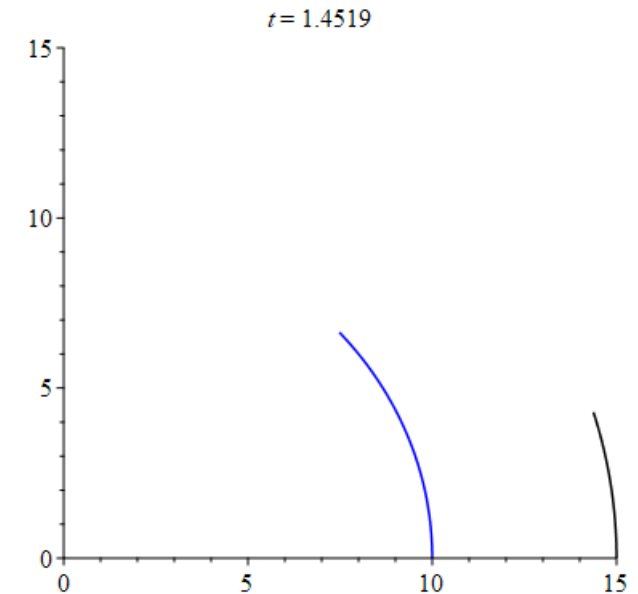
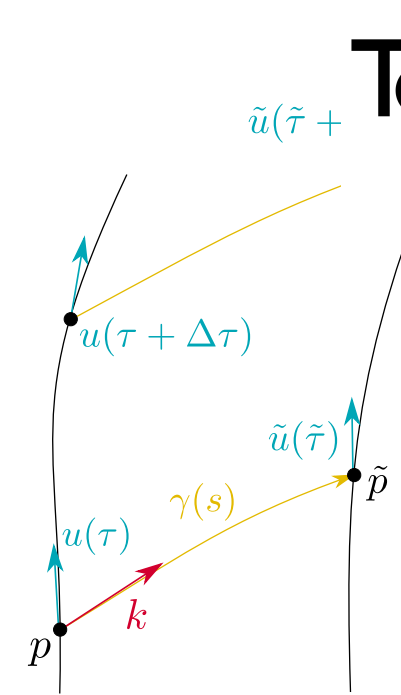
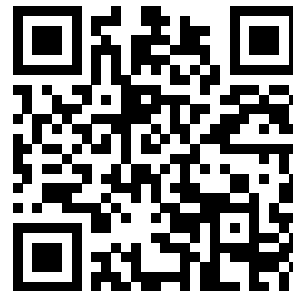
Clocks in space

Redshift calculation

- To determine redshift z we need the connecting light ray
 - General Relativistic Emitter-Observer Problem (EOP)
 - With moving boundaries

- Software GREOPy:
 - General Relativistic Emitter-Observer Python algorithm
 - Solution of EOP in arbitrary stationary spacetimes
 - Between two arbitrarily moving objects
 - So far only first order image

GREOPy link

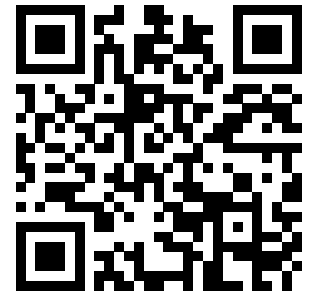


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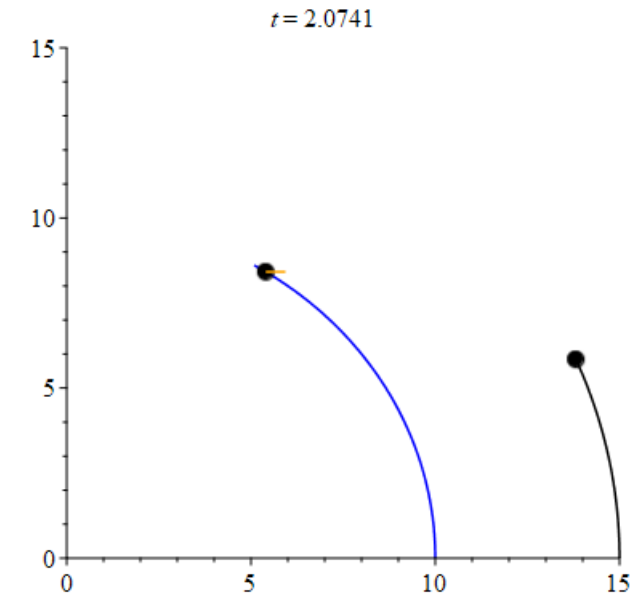
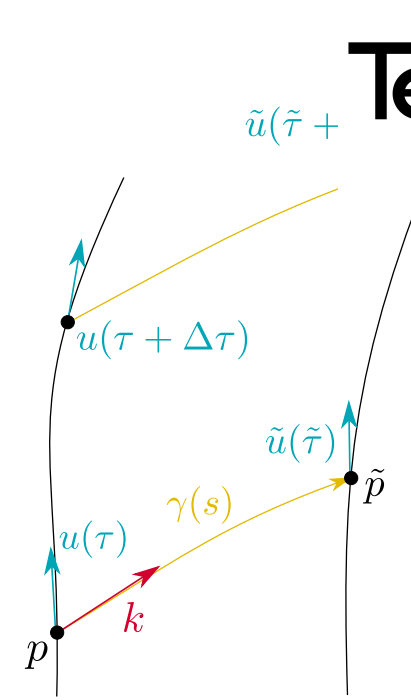
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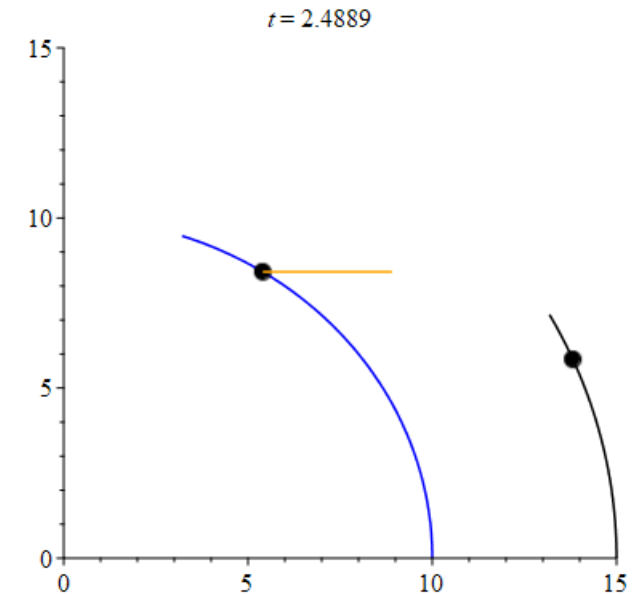
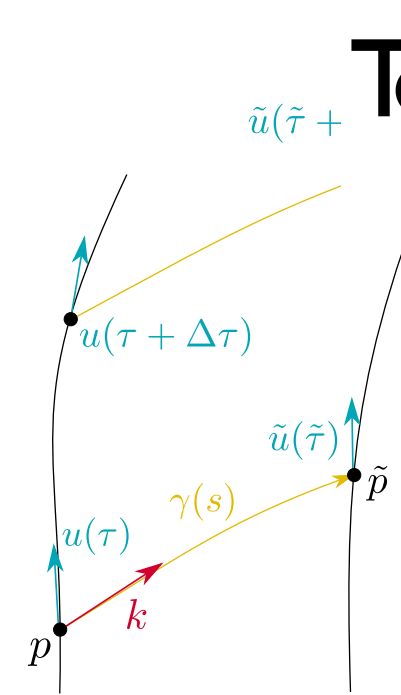
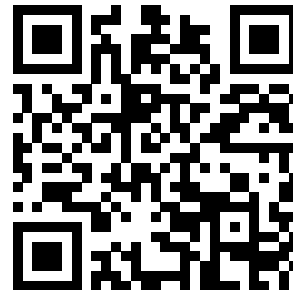
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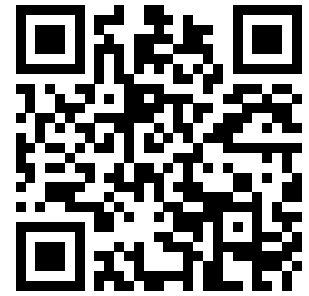


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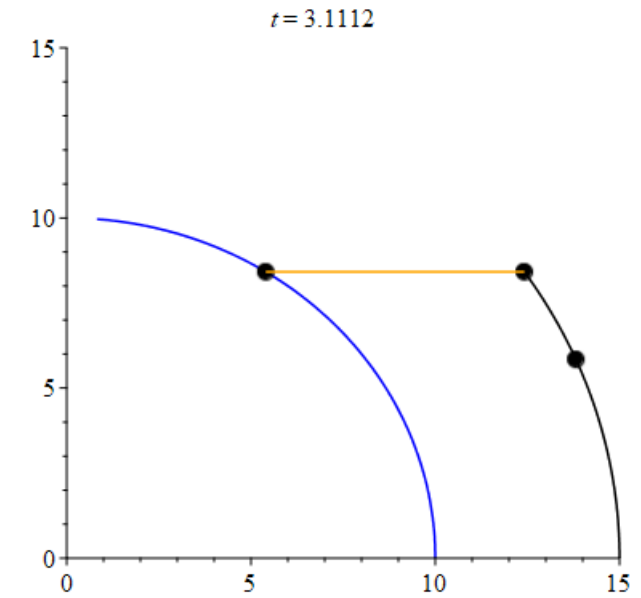
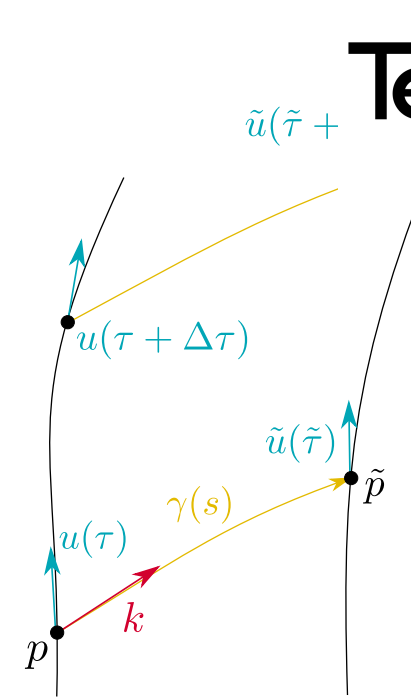
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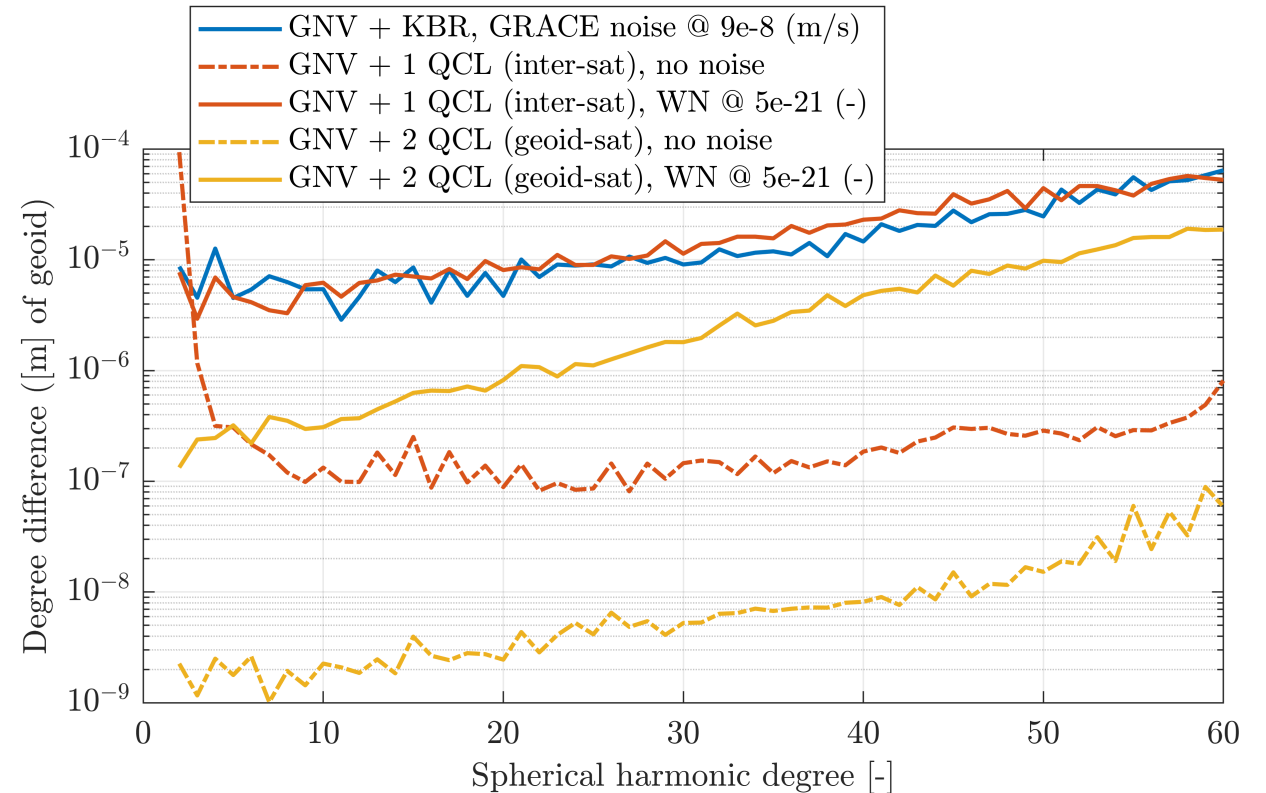
GREOPy link



Clocks in space

Preliminary study

- Closed loop simulation with XHPS
- GRACE-like mission scenario
- Assume GNSS navigation (GNV) with 2 cm white noise
- Assume K-band ranging (KBR) with typical GRACE noise
- Assume extremely precise clocks (QCL)
- Estimate spherical harmonics up to degree and order 60



Gravity Field Recovery technically possible, but...

- GRACE-setup not well suited
- Clocks needs to be extremely precise

Take Home Messages

- Gravity field observations → climate variables
- Clocks are a new tool to observe Earth's gravity field
 - Height resolution on (sub-)cm level in the next years!
- General Relativistic theoretical framework of Geodesy under development
 - Role of non-Newtonian gravitational degrees of freedom?
- Clocks in space are technically feasible for global Gravity Field Recovery, but practically not precise enough

